

# DIGITAL COMPUTER PROGRAM FOR GENERAL

MISSILE DYNAMICS SIMULATION

(N-STAGE)

SUMMARY OF EQUATIONS

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(APPENDIX B REVISED 6/20/62)

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# TABLE OF CONTENTS

TITLE	rage Number
Introduction	iv
Derivatives of Equations of Motion	1
Thrust, Mass Flow	1
Aerodynamic Drag	4
Center of Gravity Offset Force	4
Aerodynamic Normal Force	5
Resultant Forces	9
Gravity	9
Acceleration	10
Integrals for Velocity, Position and Mass	11
Turning Rate Options	11
Integrals for Attitude	15
Auxiliary Computations	16
Descriptive Quantities	16
Miss Distances	21
Orbital and Impact Approximations	25
Velocity Losses and Impulse Calculations	29
Radar Conversions	31
Radar Site Computations	31
Site Centered Cartesian	31
Spherical Radar	34
Doppler Radar	35
Site Centered Cartesian  Spherical Radar  Doppler Radar  Look Angles  Initialization  Launch Azimuth Estimation  Launch Attitude  Position and Velocity Components at Launch	38
Initialization	<b>5</b> , 44
Launch Azimuth Estimation	<b>%</b> 44
Launch Attitude	45
Position and Velocity Components at Launch	47
Radar to Cartesian Coordinate Conversion	50
Attitude from Angles of Attack	54
The parties of the state of the	56

### TABLE OF CONTENTS

Title	Page Number
Appendices	57
Appendix A - Symbol Definition	A-1,2
Appendix B - Iteration	B-1
Appendix C - Bibliography	C-1

「我、配有工作品を存在を選手をあるのか、好意、から、そのものがあます。

# LIST OF FIGURES

Pigure Number	Title	Page Number
ı	Cartesian Inertial Coordinate System with Attitude Unit Vectors	2
2	Angles of Attack	6
3	Center of Gravity	8
14	Beta.	17
5	Beta I	17
6	Epsilon Local	17
7	Epsilon	17
8	Latitude and Longitude	19
9	Down-range and Cross-range Miss Distances	24
10	Tracker Coordinates on Reference Ellipsoid	32
n	Coordinate Transformations	39
12	Rectangular to Spherical Coordinate Transformation	40
13	Doppler Rate Station Set-Up	40
14	Radar Look Angles	43
15	Initial Attitude Voctors	46
16	Launch Coordinates on Reference Ellipsoid	49

This report presents a summary of the mathematical equations contained in the N-STACE missile simulation program. This program, written for the IBM 7090 digital computer, is designed to be a flexible, high speed computation tool, useful for guidance system analysis, trajectory design, hardware evaluation, or post-flight analysis of any current or proposed missile system. The simulation is generally complete and in most instances accuracy will be limited only by the accuracy of the input data.

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Currently, work is proceeding which will add the effects of missile rotational inertia to the simulation.

The mathematical model used is fundamentally simple. The forces acting on an assumed "point mass" are computed and divided by the mass of the missile to obtain accelerations. These accelerations are added vectorially to that of gravity and numerically integrated to obtain a velocity increment over a short period of time. The velocity is numerically integrated to obtain a position increment for the same period of time. As this process is repeated, the normal output of the simulation is the time history of a trajectory.

Discontinuities in the trajectory, such as stagings, jettisons, etc., are functions of input data and are treated logically by the program.

In addition to the quantities necessary to the simulation, other descriptive quantities are computed and printed.

For design and post-flight analysis problems, automatic iterations (over trajectories or part: I trajectories) may be made by the program. The iteration techniques allow any number of input parameters to be found, with reference to values for any number of fundamental or descriptive quantities. Maximization or minimization may also be accomplished, with or without the presence of other constraints.

Although numerical integration is at the heart of the program, no statement of the integration techniques used is here made. The interested reader is referred to any standard numerical analysis text for a discussion of Runge-Kutta and Adams-Moulton integration techniques.

All of the basic computations are carried out in a Cartesian inertial system having its origin at the center of the reference ellipsoid. The x and y axis pass through the equator and the z axis is along the polar axis. (Fig. 1)

Missile attitude determines the directions of the force vectors. Mutually perpendicular unit vectors are defined such that  $\vec{\xi}$  points along the missile roll axis,  $\vec{\eta}$  lies in the pitch plane, and  $\vec{\zeta}$  lies in the yaw plane. (Fig. 1)

### Thrust and Mass Flow

are given by

$$F \xi_F = K_F F_0$$

where

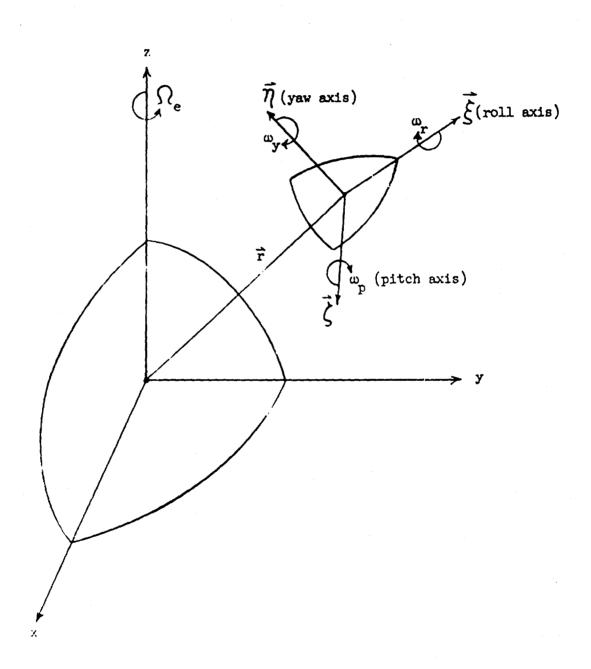
 $F_{_{\rm O}}$ ,  $\dot{M}_{_{\rm O}}$  may result from a table interrogation or from the Atlas Influence Coefficient Engine Model.

Velocity relative to air mass is

$$\vec{V}_{a} = \begin{bmatrix} \dot{x} + \Omega_{e} y - V_{wx} \\ \dot{y} - \Omega_{e} x - V_{wy} \\ \dot{z} + 0 - V_{wz} \end{bmatrix}$$

where wind velocity components are

$$V_{wx} = -W_{\phi} = \frac{xz}{\left|\overrightarrow{r}\right| \left[x^2 + y^2\right]^{1/2}} - W_{\phi} = \frac{y}{\left[x^2 + y^2\right]^{1/2}}$$



MARIESIAN INERTIAL COORDINATE SYSTEM WITH ATTITUDE UNIT VECTORS

### Thrust and Mass Flow (Cont'd)

$$v_{yy} = -w_{\phi} \frac{yz}{|\hat{r}| [x^2 + y^2]^{1/2}} + w_{\phi} \frac{x}{[x^2 + y^2]^{1/2}}$$

$$V_{W2} = W_{\phi} \frac{\left[x^2 + y^2\right]^{1/2}}{\left|\vec{r}\right|}$$

and

$$W_{\phi} = -V_{w} K_{p} \cos A_{z}$$

(north component of wind)

$$W_e = -V_w K_p sinA_z$$

(east component of wind)

Atmospheric quantities  $\rho$  (alugs/ft<sup>3</sup>), C (ft/sec), T (deg Rankine) and P (lbs/in<sup>2</sup>) are determined by approximation to the ARDC model of 1959 and are available to 2.3 million feet of altitude.

Radius at sea level

$$R_{SL} = \overline{A} \left[ 1 - \frac{1}{2} \left( k \frac{z}{\left| \overrightarrow{r} \right|} \right)^2 + \frac{3}{8} \left( k \frac{z}{\left| \overrightarrow{r} \right|} \right)^{l_t} \right]$$

where

$$k^2 = \frac{2e - e^2}{(1 - e)^2}$$

e ellipticity (flattening)

Altitude

$$n = |\hat{r}| - R_{SL}$$

# Thrust and Mass Flow (Cont'd)

Mach number

$$N_{\rm m} = \frac{|V_{\rm a}|}{c}$$

Dynamic pressure

$$Q = \frac{1}{2} \rho \left| \overrightarrow{V}_{a} \right|^{2}$$

### Aerodynamic Drag

$$F \xi_D = C_D Q S$$

where  $C_{\mathrm{D}}$  is obtained from a table, normally as a function of  $N_{\mathrm{m}}$ .

# Center of Gravity Offset Force

 $\xi_{\rm CG}$  (See Fig. 3) is obtained directly from a table or  $\overline{\xi}_{\rm CG}$  is obtained from a table and

$$\xi_{\rm CG} = L - \overline{\xi}_{\rm CG}$$

where  $\overline{\xi}_{\rm CG}$  is nose to center of gravity distance along  $\overline{\xi}$  axis. The CG distances in the  $\eta$  direction ( $\eta_{\rm CG}$ ) and  $\zeta$  direction ( $\zeta_{\rm CG}$ ) are tabular.

Center of gravity offset forces due to drag in  $\eta$  and  $\zeta$  directions are

$$F\eta_{D} = F\xi_{D} \cdot (\frac{\eta_{CG}}{\xi_{CG}})$$

$$F_{\zeta_D} = F_{\xi_D} = (\frac{\zeta_{CG}}{\xi_{CG}})$$

# Center of Gravity Offset Force (Cont'd)

Center of gravity offset forces due to thrust misalignment in  $\eta$  and  $\zeta$  directions are

$${}^{F}\eta_{F} = {}^{F}\xi_{F} \left(\frac{\eta_{CG}}{\xi_{CG}}\right)$$

$$F_{\zeta_F} = F_{\xi_F} \left( \frac{\zeta_{CG}}{\xi_{CG}} \right)$$

### Aerodynamic Normal Force

Pitch and yaw components of angle of attack ( $\alpha$ ), angle between  $\xi$  and  $\overline{V_a}$ , are determined from

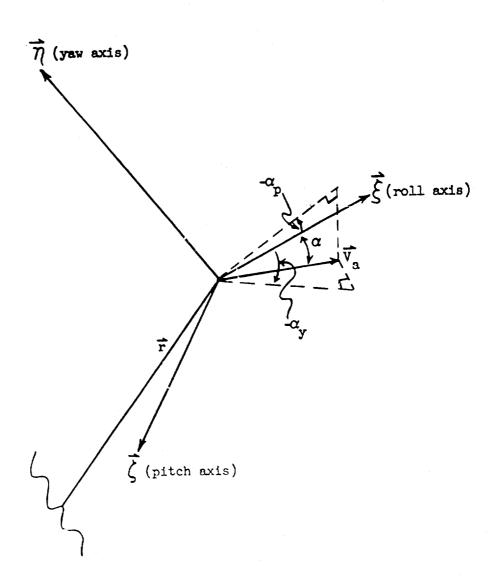
$$\tan \alpha_{\mathbf{p}} = -\frac{\overrightarrow{\eta} \cdot \overrightarrow{\mathbf{v}_{\mathbf{a}}}}{\overrightarrow{\xi} \cdot \overrightarrow{\mathbf{v}_{\mathbf{a}}}} = -\frac{\overrightarrow{\eta} \cdot \overrightarrow{\mathbf{v}_{\mathbf{a}}}}{|\overrightarrow{\mathbf{v}_{\mathbf{a}}}|}$$

(the approximate form is useful for small values of  $\alpha$ )

$$\tan \alpha_y = -\frac{\overrightarrow{\zeta} \cdot \overrightarrow{v_a}}{\overrightarrow{\xi} \cdot \overrightarrow{v_a}} = -\frac{\overrightarrow{\zeta} \cdot \overrightarrow{v_a}}{|\overrightarrow{v_a}|}$$

$$\tan \alpha = \left[ \tan^2 \alpha_p + \tan^2 \alpha_y \right]^{1/2}$$

If the approximation is used,  $\alpha_p$  and  $\alpha_y$  are assumed equal, respectively, to tan  $\alpha_p$  and tan  $\alpha_p$  .



ANGLES OF ATTACK

FIGURE 2

# Aerodynamic Normal Force (Cont'd)

Aerodynamic normal force coefficients in the pitch and yaw planes ( $C_{N_p}$ ,  $C_{N_y}$ ) may be presented as tabular functions of  $\alpha_p$ ,  $\alpha_y$  and  $N_m$ . When  $C_{N_p}$  and  $C_{N_y}$  are presented

$$C_{N_p} = C_{N_{p_\alpha}} \alpha_p$$

$$C_{N_y} = C_{N_{y_\alpha}} \alpha_y$$

or, a tabular normal force coefficient may be entered directly or as a function of total angle of attack:

$$C_N = C_N \alpha$$

$$C_{N_{p}} = \frac{\tan \alpha}{\tan \alpha} (C_{N})$$

$$C_{N_{_{_{\boldsymbol{V}}}}} = \frac{\tan \alpha_{_{_{\boldsymbol{Y}}}}}{\tan \alpha} (C_{N})$$

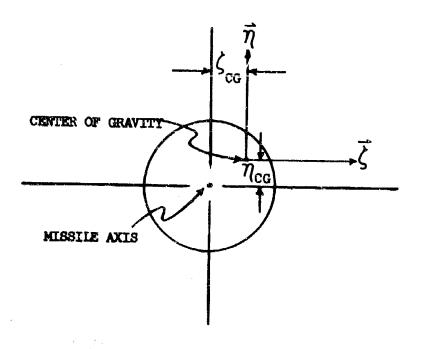
Either  $\xi_{\rm cp}\eta$ ,  $\zeta$  are obtained directly from a table or  $\overline{\xi}_{\rm cp}\eta$ ,  $\zeta$  are obtained from a table and

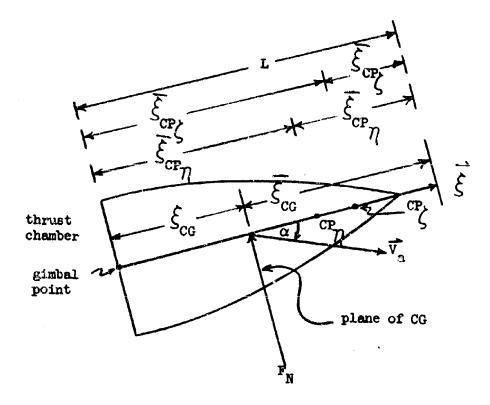
$$\xi_{\rm cp}$$
  $\eta, \zeta = L - \overline{\xi}_{\rm cp}$ 

are the nose to center of pressure distances in the pitch and yaw planes

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CENTER OF GRAVITY

FIGURE 3

# Aerodynamic Normal Force (Cont'd)

Aerodynamic normal forces in  $\overline{\eta}$  and  $\overline{\zeta}$  directions are

$$F\eta_{N} = c_{N_{p}} \frac{\xi_{cp}}{\xi_{ce}} es$$

$$F\zeta_{N} = c_{N_{y}} \frac{\xi_{cp}}{\xi_{ce}} es$$

If

$$\xi_{\rm cp}$$
 (or  $\overline{\xi}_{\rm cp}$ ) is not supplied,  $\xi_{\rm cp}\eta$  (or  $\overline{\xi}_{\rm cp}\eta$ ) is used in both planes.

# Resultant Forces

$$\xi = \xi_F - \xi_D$$

$$^{F}\eta = ^{F}\eta_{N} + ^{F}\eta_{F} - ^{F}\eta_{D}$$

$$F_{\zeta} = F_{\zeta_N} + F_{\zeta_F} - F_{\zeta_D}$$

# Gravity

$$3_{x} = \epsilon_{1} \left( \frac{x}{|r|^{3}} \right)$$

$$G_y = \epsilon_1 \left( \frac{y}{|\dot{r}|^3} \right)$$

$$G_z = \epsilon_2 \left( \frac{z}{|z|^3} \right)$$

Gravity (Cont'd)

$$\mathbf{e}_{1} = \frac{-\mathbf{G}_{\mathbf{x}}}{\left|\vec{\mathbf{r}}\right|^{2}} \left[ \left|\vec{\mathbf{r}}\right|^{2} + \mathbf{J}\vec{\mathbf{A}}^{2} \left(1 - \frac{5z^{2}}{\left|\vec{\mathbf{r}}\right|^{2}}\right) - \frac{\mathbf{H}\vec{\mathbf{A}}^{2}}{\left|\vec{\mathbf{r}}\right|^{2}} \left(-3 + \frac{7z^{2}}{\left|\vec{\mathbf{r}}\right|^{2}}\right) + \frac{\mathbf{D}\vec{\mathbf{A}}^{4}}{\left|\vec{\mathbf{r}}\right|^{2}} \left(\frac{9z^{4}}{\left|\vec{\mathbf{r}}\right|^{4}} - \frac{6z^{2}}{\left|\vec{\mathbf{r}}\right|^{2}} + \frac{3}{7}\right) \right]$$

$$\epsilon_{2} = \frac{-0_{m}}{|\vec{r}|^{2}} \left[ 2J\vec{A}^{2} - \frac{R\vec{A}^{3}Z}{|\vec{r}|^{2}} \left( -3 + \frac{3|\vec{r}|^{2}}{5Z^{2}} \right) - \frac{D\vec{A}^{4}}{|\vec{r}|^{2}} \left( \frac{4Z^{2}}{|\vec{r}|^{2}} - \frac{12}{7} \right) \right] + \epsilon_{1}$$

and J, H and D are constants.  $\epsilon_1$  and  $\epsilon_2$  optionally need not be recomputed for small changes in z position.

Acceleration

$$\vec{\dot{\mathbf{v}}} = \frac{1}{\mathbf{M}} \begin{bmatrix} \xi_{\mathbf{x}} & \eta_{\mathbf{x}} & \zeta_{\mathbf{x}} \\ \xi_{\mathbf{y}} & \eta_{\mathbf{y}} & \zeta_{\mathbf{y}} \\ \xi_{\mathbf{z}} & \eta_{\mathbf{z}} & \zeta_{\mathbf{z}} \end{bmatrix} \begin{bmatrix} \mathbf{F}_{\xi} \\ \mathbf{F}_{\eta} \\ \mathbf{F}_{\zeta} \end{bmatrix} + \vec{\mathbf{G}}$$

or  $\vec{v} = -\vec{v}_a \left[ \frac{\vec{F} \xi_D}{\vec{N} | \vec{V}_D} \right] + \vec{G}$ 

in the case of spherical body drag. In this case ablation effects may be included:

where  $R_{\underline{M}}$  and  $R_{\underline{S}}$  are tabular mass and cross-sectional area ratio factors.

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# Integrals for Velocity, Position and Mass

$$\vec{v} = \int \vec{v} dt$$

$$\vec{r} = \int \vec{v} dt$$

$$M = \int \dot{M} dt$$

### Turning Rate Options

Several computed turning rate options are available. In each option the rate computed  $(\omega_g)$  is limited to  $\pm$  2° per integration cycle.

Zero lift flight (gravity turn)

Initially orient  $\overrightarrow{V}$  to obtain  $\alpha = 0$ 

$$\vec{v}_a = |\vec{v}_a| \vec{\xi}$$

then

$$\vec{v} = \Omega_e \begin{bmatrix} -y \\ x \\ 0 \end{bmatrix} + \vec{v}_a + \vec{v}_w$$

and compute pitch and yaw rates to maintain a constant attitude relationship between  $\frac{1}{\xi}$  and  $\frac{1}{v_a}$ 

$$\omega_{pg} = -\frac{\overrightarrow{\eta} \cdot \overrightarrow{v_a}}{\overrightarrow{\xi} \cdot \overrightarrow{V_a}}$$

$$v_{yg} = \frac{\vec{\zeta} \cdot \vec{v}_{a}}{\vec{\xi} \cdot \vec{v}_{a}}$$

# Turning Rate Options (Cont'd)

where

$$\dot{V}_{ax} = \dot{V}_{x} + \frac{x}{(x^{2} + y^{2})} \left[ (V_{a})^{2} + (V_{wz})^{2} \right]$$

$$-\frac{vwz \, w_{\bullet}}{(x^2 + y^2)^{1/2}} + 2 \, \Omega_e \, v_{wy} + (\Omega_e)^2 \, x$$

$$\dot{V}_{ay} = \dot{V}_{y} + \frac{y}{(x^{2} + y^{2})} \left[ (V_{wz})^{2} + (V_{wz})^{2} \right]$$

$$+ \frac{vw1 w_{e}}{(x^{2} + y^{2})^{1/2}} - 2\Omega_{e} v_{wx} + (\Omega_{e})^{2} y$$

$$\dot{V}_{az} = \dot{V}_z + \frac{z}{\left|\frac{z}{r}\right|^2} \left(V_{\phi}\right)^2$$

and

$$VW1 = \frac{(W_f) \times z}{\left|\frac{1}{x^2}\right| \cdot (x^2 + y^2)^{1/2}}$$

$$VW2 = \frac{(W_1) \ y \ z}{\left|\vec{r}\right| \ (x^2 + y^2)^{1/2}}$$

### Turning Rate Options (Cont'd)

Roll to maintain pitch plane through center of Cartesian inertial system

$$\omega_{rg} = -\frac{\vec{v} \cdot \vec{\zeta}}{\vec{r} \cdot \vec{\eta}} + \frac{\vec{r} \cdot \vec{\xi}}{\vec{r} \cdot \vec{\eta}} \quad \omega_{y}$$

Note that  $\omega_{rg}$  is indeterminant during vertical rise.

Instantaneous turns about pitch, yaw and/or roll axes (Fig. 2).

$$\vec{\xi} = \vec{\xi}_{0} \cos \theta_{p} - \vec{\eta}_{0} \sin \theta_{p}$$

$$\vec{\eta} = \vec{\xi}_{0} \sin \theta_{p} + \vec{\eta}_{0} \cos \theta_{p}$$

$$\vec{\zeta} = \vec{\zeta}_{0}$$

$$\vec{\xi} = \vec{\xi}_{0} \cos \theta_{y} + \vec{\xi}_{0} \sin \theta_{y}$$

$$\vec{\eta} = \vec{\eta}_{0}$$

$$\vec{\zeta} = -\vec{\zeta}_{0} \sin \theta_{y} + \vec{\zeta}_{0} \cos \theta_{y}$$

$$\vec{\xi} = \vec{\xi}_{0}$$

$$\vec{\eta} = \vec{\eta}_{0} \cos \theta_{r} - \vec{\zeta}_{0} \sin \theta_{r}$$

$$\vec{\zeta} = \vec{\eta}_{0} \sin \theta_{r} + \vec{\zeta}_{0} \cos \theta_{r}$$

Turning Rate Options (Cont'd)

Obtain a specified epsilon local  $(\epsilon_{L_S})$  (see below) in a specified time  $(\Delta t)$  from the current  $\epsilon_L(\epsilon_{L_S})$ 

$$\omega_{pg} = \frac{\epsilon_{L_g} - \epsilon_{L_C}}{\Delta^t}$$

where

 $\epsilon_{L}$  is the angle between  $\xi$  and r. (Fig. 6)

Obtain  $\epsilon_{L_S}$  immediately from  $\epsilon_{L_C}$ 

$$\vec{\xi} = \vec{\xi}_{o} \cos \Psi - \vec{\eta}_{o} \sin \Psi$$

$$\vec{\eta} = \vec{\xi}_0 \sin \Psi + \vec{\eta}_0 \cos \Psi$$

Make  $\epsilon_L = \beta_I$ , angle between  $\vec{r}$  and  $\vec{v}$  (Fig. 5)

$$\omega_{pg} = \frac{\beta_{I} - \epsilon_{L_{C}}}{\delta t} + \dot{\epsilon}_{L}$$

where 8t is the computing increment and

$$\epsilon_{L} = \frac{|\vec{\mathbf{v}}| \sin \beta_{\mathbf{I}}}{|\vec{\mathbf{r}}|}$$

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# Turning Rate Options (Cont'd)

Maintain a specified  $\epsilon_{L}$ 

$$\omega_{pg} = \frac{\epsilon_{L_{M}} - \epsilon_{L_{C}}}{\delta t} + \dot{\epsilon}_{L}$$

where  $\epsilon_L$  is the  $\epsilon_L$  to be maintained.

# Integrals for Attitude

Integrands for attitude are

$$\dot{\xi} = -\hat{\eta} \omega_{p} + \hat{\zeta} \omega_{y}$$

$$\dot{\zeta} = -\hat{\xi} \omega_{y} + \hat{\eta} \omega_{r}$$

Attitudes are

$$\frac{\vec{\xi}}{\vec{\zeta}} = \int \frac{\vec{\xi}}{\vec{\zeta}} dt$$

$$\frac{\vec{\eta}}{\vec{\eta}} = \vec{\zeta} \times \vec{\xi}$$

where

### Descriptive Quantities

These equations specify some descriptive quantities used for special purposes and/or print out.

Angle between  $\vec{V}_{a}$  and  $\vec{r}$  (Fig. 4)

$$\beta = \cos^{-1} \frac{\vec{r} \cdot \vec{v}_a}{|\vec{r}| |\vec{v}_a|} \qquad 0 \le \beta \le 180^{\circ}$$

$$\Gamma = \frac{\pi}{2} - \beta$$

Angle between r and V (Fig. 5)

$$\beta_{\underline{I}} = \cos^{-1} \frac{\overrightarrow{r} \cdot \overrightarrow{\dot{r}}}{|\overrightarrow{r}| |\overrightarrow{\dot{r}}|} \qquad 0 \le \beta_{\underline{I}} \le 180^{\circ}$$

$$\int_{T} = \frac{\pi}{2} - \beta_{T}$$

Angle between  $\vec{r}$  and  $\vec{\xi}$  (Fig. 6)

$$\epsilon_{L} = \cos^{-1} \frac{\vec{r} \cdot \vec{\xi}}{|\vec{r}|}$$
  $0 \le \epsilon_{L} \le 180^{\circ}$ 

Angle between  $\hat{\xi}$  and launch vertical ( $\hat{\xi}_{L}$ ) (Fig. 7)

$$\epsilon = \cos^{-1} \quad \vec{\xi} \cdot \vec{\xi}_{L} \qquad 0 \le \epsilon \le 180^{\circ}$$

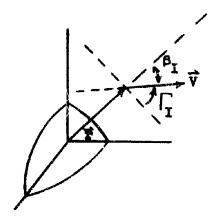
Azimuth from north of 5 projected on a plane normal to r

$$A_{\text{body}} = \tan^{-1} \frac{\left| \overrightarrow{r} \right| \left[ x \ \xi_{y} - y \ \xi_{x} \right]}{y \left[ y \ \xi_{z} - z \ \xi_{y} \right] - x \left[ z \ \xi_{x} - x \ \xi_{z} \right]}$$

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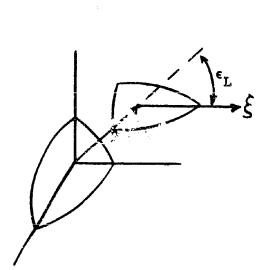
BETA

FIGURE 4



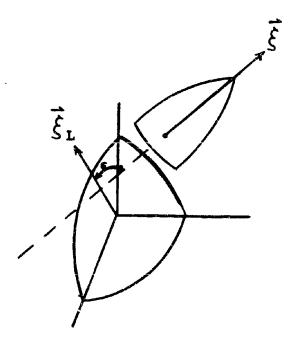
BETI

FIGURE 5



EPSILON LOCAL

FIGURE 6



EPSILON

7

FIGURE

# Descri tive Quantities (Cont'd)

Normalized angular momentum

RVSB = 
$$|\vec{r}| |\vec{v}| \sin \beta_I$$

Ratio squared of velocity to circular orbital velocity

RV8G = 
$$\frac{|\vec{\mathbf{y}}|^2}{G_m}$$

Total energy

ENGY = 
$$\frac{|\vec{v}|^2}{2} - \frac{G_m}{|\vec{r}|}$$

Vis viva energy

Sensed velocity

$$V_{S} = \int (\vec{\dot{V}} - \vec{G}) dt$$

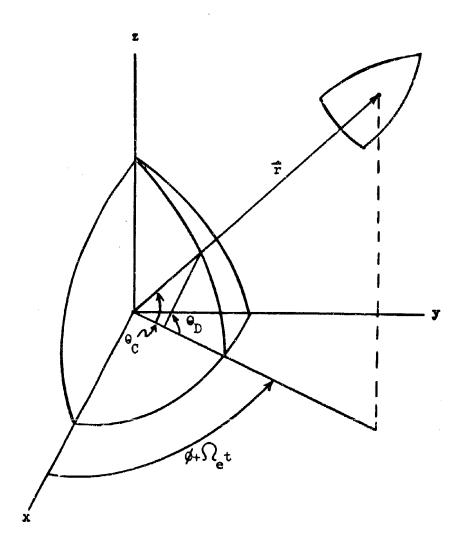
Aerodynamic heating

$$H_A = \int QV_a dt$$

Geodetic latitude ( $\theta_{\rm D}$ ) and longitude ( $\emptyset$ )(Fig. 8)

$$e_D = tan^{-1} \frac{z}{(1-e)^2(x^2+y^2)^{1/2}}$$

$$\emptyset = \left[ \tan^{-1} \left( \frac{y}{x} \right) \right] - \Omega_{e} t$$



LATITUDE / LONGITUDE

FIGURE 8

Descriptive Quantities (Cont'd)

 $\theta_{c} = \tan^{-1} \left[ \frac{z}{(z^2 + z^2)^{1/2}} \right]$ 

Circular range (R) between two positions on the spheroid, where  $(\theta_{c1}, \phi_1)$  (geocentric latitude, longitude) defines the first position,  $(\theta_{c2}, \phi_2)$  defines the second, and  $r_0$  is the radius of the spheroid, is given by

$$\mathbf{E}_{\mathbf{Y}} = \sin \theta_{c1} \sin \theta_{c2} + \cos \theta_{c1} \cos \theta_{c2} \cos (\phi_{2} - \phi_{1})$$

If

$$\mathbf{E}_{\mathbf{\Psi}} \geq .99998$$
, then

$$\Psi = \left[ (\theta_{c2} - \theta_{c1})^2 + (\phi_2 - \phi_1)^2 \cos^2 \theta_{c1} \right]^{1/2}$$

where

$$|\theta_{c2} - \theta_{c1}|$$
 and  $|\phi_2 - \phi_1|$  are  $\leq 180^\circ$ 

If

$$\mathbf{E}_{\mathbf{V}}$$
 < .99998, then

$$\Psi = \cos^{-1} \mathbf{E} \mathbf{\Psi}$$

$$R = \begin{cases} 2\pi - \Psi, & \text{if } 90^{\circ} \le |p| \le 180^{\circ} \\ \Psi, & \text{if } 0^{\circ} \le |p| < 90^{\circ} \end{cases}$$

where p is the clockwise difference of the azimuth from  $(\Theta_{c1}, \phi_1)$  to  $(\Theta_{c2}, \phi_2)$  and a reference azimuth.

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Miss

tote

### Descriptive Quantities (Cont'd)

Azimuth from North of V projected on a plane normal to r is

$$A_{V} = \tan^{-1} \left[ \frac{\dot{y} \cos(\Omega_{e} t + \phi) - \dot{x} \sin(\Omega_{e} t + \phi)}{\dot{z} \cos(\Omega_{e} - \sin(\Omega_{e} t + \phi) + \dot{y} \sin(\Omega_{e} t + \phi)} \right]$$

Cartesian inertial to launch centered inertial coordinate transformation

$$\begin{bmatrix} \mathbf{p}_{xg} \\ \mathbf{p}_{yg} \\ \mathbf{p}_{zg} \end{bmatrix} = \begin{bmatrix} -\eta_{x_0} & -\eta_{y_0} & -\eta_{z_0} \\ -\zeta_{x_0} & -\zeta_{y_0} & -\zeta_{z_0} \\ -\xi_{x_0} & -\xi_{y_0} & -\xi_{z_0} \end{bmatrix} \begin{bmatrix} \mathbf{x} \\ \mathbf{y} \\ \mathbf{z} \end{bmatrix} - \begin{bmatrix} -\eta_{x_0} & -\eta_{y_0} & -\eta_{z_0} \\ -\zeta_{x_0} & -\zeta_{y_0} & -\zeta_{z_0} \\ -\xi_{x_0} & -\xi_{y_0} & -\xi_{z_0} \end{bmatrix} \begin{bmatrix} \mathbf{x} \\ \mathbf{y} \\ \mathbf{z} \end{bmatrix}$$

$$\begin{bmatrix} v_{xg} \\ v_{yg} \\ v_{zg} \end{bmatrix} = \begin{bmatrix} -\eta_{x_0} & -\eta_{y_0} & -\eta_{z_0} \\ -\zeta_{x_0} & -\zeta_{y_0} & -\zeta_{z_0} \\ -\xi_{x_0} & -\xi_{y_0} & -\xi_{z_0} \end{bmatrix} \begin{bmatrix} \dot{x} \\ \dot{y} \\ \dot{z} \end{bmatrix}$$

where  $\overline{\xi}_{o}$ ,  $\overline{\eta}_{o}$ ,  $\overline{\zeta}_{o}$  are the attitude vectors at launch.

#### Miss Distances

Cross range miss distance  $(M_c)$  and down range miss distance  $(M_d)$  components of total miss distance  $(M_+)$  at impact (Fig. 9) are given by

$$E_{L,\tau} = \left[ \sin\theta_{CL} \sin\theta_{CI} + \cos\theta_{CL} \cos\theta_{CI} \cos(\phi_1 - \phi_L) \right]$$

Ιſ

$$B_{L,I} \ge .99998$$
, then

$$\Psi_{L,I} = \left[ (e_{CI} - e_{CL})^2 + (\phi_I - \phi_L)^2 \cos^2 e_{CL} \right]^{1/2}$$

where

$$|\theta_{\text{CI}} - \theta_{\text{CL}}|$$
 and  $|\phi_{\text{I}} - \phi_{\text{L}}|$  are  $\leq 180^{\circ}$ 

If

$$E_{L,I} < .99998$$
, then

$$\Psi_{\rm L,I} = \cos^{-1} E_{\rm L,I}; o \le \Psi_{\rm L,I} \le 180^{\circ}$$

$$E_{L,M} = \sin\theta_{CL} \sin\theta_{CM} + \cos\theta_{CL} \cos\theta_{CM} \cos(\phi_M - \phi_L)$$

If

$$E_{L,M} \ge .99998$$
, then

$$\Psi_{L,M} = \left[ (e_{CM} - e_{CL})^2 + (\phi_{M} - \phi_{L})^2 \cos^2 e_{CL} \right]^{1/2}$$

where

$$| \Theta_{\rm CM} - \Theta_{\rm CL} |$$
 and  $| \phi_{\rm M} - \phi_{\rm L} |$  are  $\leq 180^{\circ}$ 

If

$$E_{L,M} \leq .99998$$
, then

$$\Psi_{L,M} = \cos^{-1} E_{L,M}; 0 \le \Psi_{L,M} \le 180^{\circ}$$

# Miss Distances (Cont'd)

$$\sin \eta_{L,I} = \frac{\cos \varphi_{CI} \sin(\phi_{I} - \phi_{L})}{\sin \Psi_{L,I}}$$

$$\cos \eta_{L,I} = \frac{(\sin \theta_{CL} \cos \Psi_{L,I}) - \sin \theta_{CI}}{\cos \theta_{CL} \sin \Psi_{L,I}}$$

$$\sin 7_{L,M} = \frac{\cos \theta_{CM} \sin (\phi_{M} - \phi_{L})}{\sin L_{L,M}}$$

$$\cos \eta_{L,M} = \frac{(\sin \theta_{CL} \cos \Psi_{L,M}) - \sin \theta_{CM}}{\cos \theta_{CL} \sin \Psi_{L,M}}$$

$$\operatorname{sinM}_{\mathbf{c}}^{!} = (\operatorname{sin} \mathcal{N}_{\mathbf{L},\mathbf{M}} \operatorname{cos} \mathcal{N}_{\mathbf{L},\mathbf{I}} - \operatorname{cos} \mathcal{N}_{\mathbf{L},\mathbf{M}} \operatorname{sin} \mathcal{N}_{\mathbf{L},\mathbf{I}}) \operatorname{sin} \mathcal{Y}_{\mathbf{L},\mathbf{M}}$$

$$\cos \mathbf{N_c^i} = \left[1 - \sin^2 \mathbf{N_c^i}\right]^{1/2}$$

$$(\Psi_{L,I} + M_0^*) = \cos^{-1}\left[\frac{\cos\Psi_{L,M}}{\cosh^2_c}\right]$$

$$\mathbf{H}_{\mathbf{d}} = \mathbf{r}_{\mathbf{o}} \left[ (\Psi_{\mathbf{L},\mathbf{I}} + \mathbf{K}_{\mathbf{d}}^{i}) - \Psi_{\mathbf{L},\mathbf{I}} \right]$$

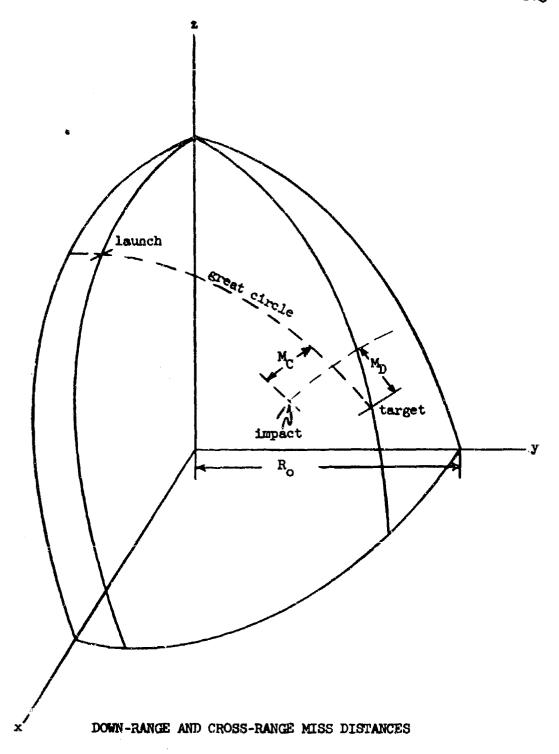


FIGURE 9

Mis

<u> Orb:</u>

# Miss Distances (Cont'd)

$$M_c = r_o M_c'$$

$$M_{t} = \left[M_{c}^{2} + M_{d}^{2}\right]^{1/2}$$

where  $M_C^i$ ,  $M_d^i$  are central angles. Note that  $M_C$  is positive to the left when looking down range.

# Orbital and Impact Approximations

For an elliptical orbit or an impact trajectory,

$$\epsilon \cos \overline{E} = \frac{|\vec{r}| \cdot |\vec{v}|^2}{G_M} - 1$$

$$\bar{A} = \frac{|\vec{r}|}{1 - \epsilon \cos \bar{k}}$$

$$\epsilon \sin \overline{E} = \frac{\overline{r} \cdot \overline{V}}{\sqrt{\overline{A} \cdot G_{M}}}$$

$$\epsilon = \left[ \left( \epsilon \cos \overline{E} \right)^2 + \left( \epsilon \sin \overline{E} \right)^2 \right]^{1/2}$$

Perigee = 
$$\overline{A}(1 - \epsilon) - r_0$$

Apogee = 
$$\overline{A}(1 + \epsilon) - r_0$$

where  $\varepsilon$  is the ellipsoidal eccentricity and  $\overline{E}$  is the eccentric anomaly of the present missile position.

# Orbital and Impact Approximations (Cont'd)

Period = 
$$\frac{2\pi \bar{A}^{3/2}}{60\sqrt{\bar{G}_{N}}}$$

Inclination =  $\cos^{-1} (\sin A_{V} \cos \theta_{CM})$ 

Velocity at apogee = 
$$\frac{\left|\vec{r}\right| \left|\vec{V}\right| \left(1 - \cos^2 \beta_{\rm I}\right)}{\bar{\Lambda}(1 + \epsilon)}$$

For an impact trajectory  $R_{SEI} \geq \overline{A}(1 - e)$ 

$$\epsilon \cos \mathbf{E}_{\mathrm{T}} = 1 - \frac{\mathbf{R}_{\mathrm{SEI}}}{\overline{\mathbf{A}}}$$

$$\epsilon \sin \mathbf{E}_{\mathrm{T}} = -\left[\epsilon^2 - (\epsilon \cos \mathbf{E}_{\mathrm{T}})^2\right]^{1/2}$$

$$\sin(\mathbf{E}_{\mathrm{T}} - \bar{\mathbf{E}}) = \frac{(\epsilon \sin \mathbf{E}_{\mathrm{T}})(\epsilon \cos \bar{\mathbf{E}}) - (\epsilon \cos \mathbf{E}_{\mathrm{T}})(\epsilon \sin \bar{\mathbf{E}})}{\epsilon^{2}}$$

$$\cos(\mathbf{E}_{\mathrm{T}} - \mathbf{E}) = \frac{(\epsilon \cos \mathbf{E}_{\mathrm{T}})(\epsilon \cos \mathbf{E}) + (\epsilon \sin \mathbf{E}_{\mathrm{T}})(\epsilon \sin \mathbf{E})}{\epsilon^{2}}$$

$$\mathbf{E}_{\mathrm{T}} \sim \mathbf{\overline{E}} = \tan^{-1} \left[ \frac{\sin(\mathbf{E}_{\mathrm{T}} - \mathbf{\overline{E}})}{\cos(\mathbf{E}_{\mathrm{T}} - \mathbf{\overline{E}})} \right]$$

where  $\boldsymbol{E}_{\boldsymbol{T}}$  is the eccentric anomaly at the impact point.

# Orbital and Impact Approximations (Cont'd)

Time of flight to impact and time of impact are

$$T_{\overline{F}} = \frac{\overline{A}^{3/2}}{\sqrt{G_{\overline{M}}}} \left[ R_{\overline{T}} - \overline{E} + \epsilon \sin \overline{E} - \epsilon \sin R_{\overline{T}} \right]$$

$$T_{\text{TMP}} = t + T_{\text{F}}$$

Inertial Cartesian components of impact point are

$$x_{ggt} = xt + xg$$

$$y_{SEI} = \dot{y}f + yg$$

$$z_{SET} = \dot{z}f + zg$$

where

$$f = \frac{\overline{A} \frac{3/2}{\sqrt{G_M}} \left[ \sin(E_T - \overline{E}) + \epsilon \sin \overline{E} - \epsilon \sin E_T \right]$$

$$g = \frac{\cos(\mathbf{E}_{\mathbf{T}^1} - \mathbf{E}) - \epsilon \cos \mathbf{E}}{1 - \epsilon \cos \mathbf{E}}$$

Azimuth of impact is

$$A_{z_{I}} = -\left[\tan^{-1} \frac{\sin(180 - A_{z_{I}})}{\cos(180 - A_{z_{I}})} - 160\right]$$

### Orbital and Impact Approximations (Cont'd)

where

$$\sin (180 - A_{21}) = \frac{\cos \theta_{C1} \sin (\phi_{1} - \phi_{L})}{\sin \Psi}$$

$$\cos (180 - A_{2I}) = \frac{\sin \theta_{CL} \cos \Psi - \sin \theta_{CI}}{\cos \theta_{CL} \sin \Psi}$$

$$E_{\Psi} = \sin\theta_{CL} \sin\theta_{CI} + \cos\theta_{CL} \cos\theta_{CI} \cos(\phi_{L} - \phi_{I})$$

Ιſ

$$\mathbf{E}_{\Psi} \ge .99998$$
, then  $\Psi = \left[ (\mathbf{e}_{cI} - \mathbf{e}_{cL})^2 + (\mathbf{e}_{I} - \mathbf{e}_{L})^2 \cos^2 \mathbf{e}_{cL} \right]^{1/2}$ 

where

$$\left| \phi_{\text{CI}} - \phi_{\text{CL}} \right|$$
 and  $\left| \phi_{\text{I}} - \phi_{\text{L}} \right|$  are  $\leq 180^{\circ}$ 

Ιť

$$E_{\Psi} = .99998, \Psi = \cos^{-1}E_{\Psi}, \phi \leq \Psi \leq 180^{\circ}$$

Circular range to impact is

$$2\pi - \Psi$$
; if  $90^{\circ} \le |p| \le 180^{\circ}$   
CRIP =  $\Psi$ ; if  $0^{\circ} \le |p| = 90^{\circ}$ 

where p is the clockwise difference of impact azimuth and launch azimuth.

# Orbital and Impact Approximations (Cont'd)

If | | - h sel or if the missile is in a hyperbolic orbit (i.e.

 $\frac{|\vec{v}|^2}{G_M} = 2$ ), all orbital and impact quantities are set to zero.

 $R_{SEI}$  (radius of spherical earth impact) may be imput directly or as a height relative to the radius of the target  $\left[R_{SEI} = f(\phi_{T}, \phi_{T}, H_{T}, EFHT)\right]$  or as a height relative to the average radius of the geoid  $(r_{O})$ .

### Velocity Losses and Impulse Calculations

Gravity loss

$$V_g = \int_{t_1}^{t_2} g \cos \beta_I dt$$

where

t, = time at beginning of stage

t, = time at end of stage

g = local gravity

 $\beta_{\tau}$  = angle between  $\vec{r}$  and geocentric local vertical  $(\vec{r})$ 

Drag loss

$$V_{d} = \int_{t_{1}}^{t_{2}} \frac{\xi}{M} \underline{D} dt$$

Misalignment loss

$$V_{m} = \int_{t_{1}}^{t_{2}} \frac{F\xi_{F}}{M} (1 - \cos \alpha) dt$$

# Velocity Losses and Impulse Calculations (Cont'd)

Average specific impulse per stage

$$\bar{\mathbf{I}}_{ap} = -\left[\frac{1}{t_2 - t_1}\right] \int_{t_1}^{t_2} \frac{\mathbf{F}_{\mathbf{F}}}{\hat{\mathbf{N}}} dt$$

Total inculse

$$I_{T} = \int_{t_{1}}^{t_{2}} F_{\xi} f dt$$

Ideal velocity per stage

$$v_{I} = g_{O} \overline{I}_{sp} \mathcal{L}_{n} (\frac{wt_{1}}{wt_{2}})$$

where  $g_0$  is the nominal value of gravity,  $wt_1$  is the weight of missile at  $t_1$ , and  $wt_2$  is the weight of missile at  $t_2$ .

### Radur Site Computations

Distance from radar site to polar axis is

$$x_{ER} = e^{-c}$$
  $\left[ r_{SL_R} \cos \theta_{CR} + (h_R + \epsilon_{SR}) \cos \theta_{DR} \right]$ 

where

$$r_{SL_{R}} \stackrel{\cos \varphi_{CR}}{=} \frac{\bar{A}}{\left[1 + (1-e)^{2} \tan^{2} \varphi_{DR}\right]^{1/2}}$$

and

h, = altitude of radar site above sea level

 $g_{ggg}$  = geoidal separation at radar site

r<sub>SI<sub>b</sub></sub> = sea level radius at radar site

 $\delta \phi_{\rm R}$  = longitude of radar, positive east

Distance of radar site above equatorial plane is

$$z_{ER} = r_{SL_R} \sin \theta_{CR} + (h_R + g_{SR}) \sin \theta_{DR}$$

where

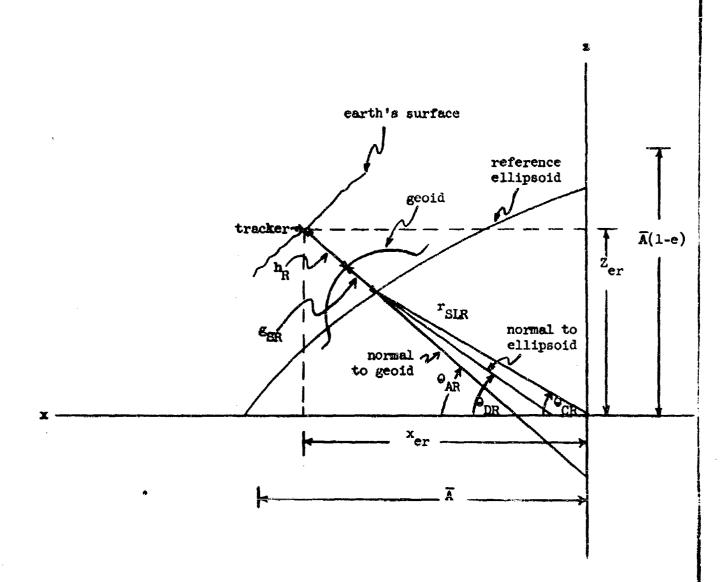
$$r_{SL_R} = in\theta_{CR} = (r_{SL_R} \cos\theta_{CR})(1-e)^2 \tan\theta_{DR}$$

# Site Centered Cartesian

Ellipsoid centered cartesian inertial coordinates converted to ellipsoidfixed rectangular with origin at redar (Fig. 10) are given by the following:

$$x = x \cos \Psi + y \sin \Psi$$

$$y_x = -x \sin \Psi + y \cos \Psi$$



TRACKER COORDINATES ON REFERENCE ELLIPSOID

FIGURE 10

# Site Centered Cartesian (Cont'd)

where

$$Y = \delta \phi_{R} + \Omega_{e} t$$

$$\overline{x}_{e} = (x_{e} - x_{e_{R}}) \sin \theta_{AR} - (z_{e} - z_{e_{R}}) \cos \theta_{AR}$$

$$\overline{y}_{e} = y_{e}$$

$$\overline{z}_{e} = (x_{e} - x_{e_{R}}) \cos \theta_{AR} + (z_{e} - z_{e_{R}}) \sin \theta_{AR}$$

$$\overline{u} = \overline{x}_{e}$$

$$\overline{v} = \overline{y}_{e} + \delta_{w_{R}} \cdot \overline{z}_{e}$$

$$\overline{v} = -\delta_{w_{R}} \cdot \overline{y}_{e} + \overline{z}_{e}$$

where  $\delta_W = \text{deviation from local vertical}$ , and  $\sin \delta_W = \delta_W$   $u = \overline{u} \cos \mu_R + \overline{v} \sin \mu_R$   $v = -\overline{u} \sin \mu_R + \overline{v} \cos \mu_R$ 

where  $\mu_{\widetilde{K}}$  = angle through which the  $\widetilde{u}$  -  $\widetilde{v}$  plane is rotated to fix reder azimuth reference

$$\dot{x}_e = \dot{x} \cos Y + \dot{y} \sin Y + \Omega_e \cdot y_e$$

$$\dot{y}_e = -\dot{x} \sin Y + \dot{y} \cos Y - \Omega_e \cdot x_e$$

$$\dot{z}_e = \dot{z}$$

# Site Centered Cartesian (Cont'd)

$$\ddot{x}_{e} = \dot{x}_{e} \sin \theta_{AR} - \dot{z}_{e} \cos \theta_{AR}$$

$$\bar{\dot{y}}_e = \dot{y}_e$$

$$\frac{1}{\hat{z}_e} = \dot{x}_e \cos\theta_{AR} + \dot{z}_e \sin\theta_{AR}$$

$$\overline{\dot{\mathbf{v}}} = \overline{\dot{\mathbf{y}}}_{\mathbf{c}} + \mathbf{b}_{\mathbf{v}} \cdot \overline{\dot{\mathbf{z}}}_{\mathbf{c}}$$

$$\dot{\bar{\mathbf{v}}} = -\mathbf{\delta}_{\mathbf{w}} \cdot \dot{\bar{\mathbf{y}}}_{\mathbf{e}} + \dot{\bar{\mathbf{z}}}_{\mathbf{e}}$$

$$\dot{\hat{u}} = \bar{\hat{v}} \cos \mu_{R} + \bar{\hat{v}} \sin \mu_{R}$$

$$\dot{\mathbf{v}} = -\dot{\mathbf{u}} \sin \mu_{\mathbf{R}} + \dot{\mathbf{v}} \cos \mu_{\mathbf{R}}$$

# Spherical Reder

Ellipsoid-fixed rectangular coordinates converted to spherical (radar) coordinates (Fig. 12)

$$R = \left[u^2 + v^2 + v^2\right]^{1/2}$$

P

Do

CO

### Spherical Redar (Cont'd)

$$A = \sin^{-1} \left[ \frac{u}{(u^2 + v^2)^{1/2}} \right] = \tan^{-1} \left( \frac{u}{v} \right), \quad 0 \le A \le 2\pi$$

$$E = \cos^{-1}\left(\frac{w}{R}\right), \quad 0 \le E \le \pi$$

$$\dot{R} = \frac{1}{R} \left( u\dot{u} + v\dot{v} + w\dot{w} \right)$$

$$\dot{A} = \frac{v\dot{u} - u\dot{v}}{(u^2 + v^2)}$$

$$\dot{E} = \frac{v(u\dot{u} + v\dot{v}) - \dot{v}(u^2 + v^2)}{R^2(u^2 + v^2)^{1/2}}$$

### Doppler Radar

Radar cross-shaped doppler rates computed from ellipsoid-fixed rectangular coordinates (Fig. 13)

$$\dot{p} = \frac{\dot{u}}{R} \left\{ u \left[ (\delta_1 - \delta_3) + \frac{3}{2} (\delta_1^2 - \delta_3^2) + \frac{5}{2} (\delta_1^3 - \delta_3^3) \right] + u_3 \left[ 1 + \delta_3 + \frac{3}{2} \delta_3^2 + \frac{5}{2} \delta_3^3 \right] - u_1 \left[ 1 + \delta_1 + \frac{3}{2} \delta_1^2 + \frac{5}{2} \delta_1^3 \right] \right\}$$

### Doppler Radar (Cont'd)

$$\begin{split} &+\frac{\dot{v}}{R}\left[v\left[\left(\delta_{1}-\delta_{3}\right)+\frac{3}{2}\left(\delta_{1}^{2}-\delta_{3}^{2}\right)+\frac{5}{2}\left(\delta_{1}^{3}-\delta_{3}^{3}\right)\right]\\ &=v_{3}\left[1+\delta_{3}+\frac{3}{2}\delta_{3}^{2}+\frac{5}{2}\delta_{3}^{3}\right]-v_{1}\left[1+\delta_{1}+\frac{3}{2}\delta_{1}^{2}+\frac{5}{2}\delta_{1}^{3}\right]\\ &+\frac{\dot{u}}{R}\left[v\left[\left(\delta_{1}-\delta_{3}\right)+\frac{3}{2}\left(\delta_{1}^{2}-\delta_{3}^{2}\right)+\frac{5}{2}\left(\delta_{1}^{3}-\delta_{3}^{3}\right)\right]\\ &+v_{3}\left[1+\delta_{3}+\frac{3}{2}\delta_{3}^{2}+\frac{5}{2}\delta_{3}^{3}\right]-v_{1}\left[1+\delta_{1}+\frac{3}{2}\delta_{1}^{2}+\frac{5}{2}\delta_{1}^{3}\right]\\ &+\frac{\dot{u}}{R}\left[u\left[\left(\delta_{2}-\delta_{4}\right)+\frac{3}{2}\left(\delta_{2}^{2}-\delta_{4}^{2}\right)+\frac{5}{2}\left(\delta_{2}^{3}-\delta_{4}^{3}\right)\right]\\ &+u_{4}\left[1+\delta_{4}+\frac{3}{2}\delta_{4}^{2}+\frac{5}{2}\delta_{4}^{3}\right]-u_{2}\left[1+\delta_{2}+\frac{3}{2}\delta_{2}^{2}+\frac{5}{2}\delta_{2}^{3}\right]\\ &+\frac{\dot{v}}{R}\left[v\left[\left(\delta_{2}-\delta_{4}\right)+\frac{3}{2}\left(\delta_{2}^{2}-\delta_{4}^{2}\right)+\frac{5}{2}\left(\delta_{2}^{3}-\delta_{4}^{3}\right)\right]\\ &+v_{4}\left[1+\delta_{4}+\frac{3}{2}\delta_{4}^{2}+\frac{5}{2}\delta_{4}^{3}\right]-v_{2}\left[1+\delta_{2}+\frac{3}{2}\delta_{2}^{2}+\frac{5}{2}\delta_{2}^{3}\right]\\ &+v_{4}\left[1+\delta_{4}+\frac{3}{2}\delta_{4}^{2}+\frac{5}{2}\delta_{4}^{3}\right]-v_{2}\left[1+\delta_{2}+\frac{3}{2}\delta_{2}^{2}+\frac{5}{2}\delta_{2}^{3}\right]\\ &+v_{4}\left[1+\delta_{4}+\frac{3}{2}\delta_{4}^{2}+\frac{5}{2}\delta_{4}^{3}\right]-v_{2}\left[1+\delta_{2}+\frac{3}{2}\delta_{2}^{2}+\frac{5}{2}\delta_{2}^{3}\right]\\ &+v_{4}\left[1+\delta_{4}+\frac{3}{2}\delta_{4}^{2}+\frac{5}{2}\delta_{4}^{3}\right]-v_{2}\left[1+\delta_{2}+\frac{3}{2}\delta_{2}^{2}+\frac{5}{2}\delta_{2}^{3}\right]\\ &+v_{4}\left[1+\delta_{4}+\frac{3}{2}\delta_{4}^{2}+\frac{5}{2}\delta_{4}^{3}\right]-v_{2}\left[1+\delta_{2}+\frac{3}{2}\delta_{2}^{2}+\frac{5}{2}\delta_{2}^{3}\right]\\ &+v_{4}\left[1+\delta_{4}+\frac{3}{2}\delta_{4}^{2}+\frac{5}{2}\delta_{4}^{3}\right]-v_{2}\left[1+\delta_{2}+\frac{3}{2}\delta_{2}^{2}+\frac{5}{2}\delta_{2}^{3}\right]\\ &+v_{5}\left[1+\delta_{4}+\frac{3}{2}\delta_{4}^{2}+\frac{5}{2}\delta_{4}^{3}\right]-v_{5}\left[1+\delta_{2}+\frac{3}{2}\delta_{2}^{2}+\frac{5}{2}\delta_{2}^{3}\right]\\ &+v_{5}\left[1+\delta_{4}+\frac{3}{2}\delta_{4}^{2}+\frac{5}{2}\delta_{4}^{3}\right]-v_{5}\left[1+\delta_{4}+\frac{3}{2}\delta_{4}^{2}+\frac{5}{2}\delta_{2}^{3}\right]\\ &+v_{5}\left[1+\delta_{5}+\frac{3}{2}\delta_{5}^{2}+\frac{5}{2}\delta_{5}^{3}\right]-v_{5}\left[1+\delta_{5}+\frac{3}{2}\delta_{5}^{2}+\frac{5}{2}\delta_{5}^{3}\right]\\ &+v_{5}\left[1+\delta_{5}+\frac{3}{2}\delta_{5}^{2}+\frac{5}{2}\delta_{5}^{3}\right]-v_{5}\left[1+\delta_{5}+\frac{3}{2}\delta_{5}^{2}+\frac{5}{2}\delta_{5}^{3}\right]\\ &+v_{5}\left[1+\delta_{5}+\frac{3}{2}\delta_{5}^{2}+\frac{5}{2}\delta_{5}^{3}\right]-v_{5}\left[1+\delta_{5}+\frac{3}{2}\delta_{5}^{2}+\frac{5}{2}\delta_{5}^{3}\right]\\ &+v_{5}\left[1+\delta_{5}+\frac{3}{2}\delta_{5}^{2}+\frac{5}{2}\delta_{5}^{3}\right]-v_{5}\left[1+\delta_{5}^{2}+\frac{3}{2}\delta_{5}^{2}+\frac{5}{2}\delta_{5}^{3}\right]\\ &+v_{5}\left[1+\delta_{5}+\frac{3}{2}\delta_{5}^{2}+\frac{5}{2}\delta_{5}^{3}\right]\\ &+v_{5}\left[1+\delta_{5}+$$

Doppler Radar (Cont'd)

$$+\frac{\dot{w}}{R}\left[w\left[(\delta_{2}-\delta_{l_{1}})+\frac{3}{2}(\delta_{2}^{2}-\delta_{l_{1}}^{2})+\frac{5}{2}(\delta_{2}^{3}-\delta_{l_{1}}^{3})\right]$$

$$+ w_{4} \left[ 1 + \delta_{4} + \frac{3}{2} \delta_{4}^{2} + \frac{5}{2} \delta_{4}^{3} \right] - w_{2} \left[ 1 + \delta_{2} + \frac{3}{2} \delta_{2}^{2} + \frac{5}{2} \delta_{2}^{3} \right]$$

where

$$\delta_{1} = \frac{(uu_{1} + vv_{1} + ww_{1}) - \frac{1}{2}(u_{1}^{2} + v_{1}^{2} + w_{1}^{2})}{R^{2}}$$

Redar L-shaped doppler position and rates computed from ellipsoid-fixed coordinates (Fig. 13) are given by the following. Constants are delineated on Page 52.

$$R^{2} = u^{2} + v^{2} + w^{2}$$

$$R_{0} = \left[R^{2} + K_{4} - uK_{1} - vK_{2} - wK_{3}\right]^{1/2}$$

$$R_{1} = \left[R^{2} + K_{6} - uK_{5} - vK_{6} - wK_{7}\right]^{1/2}$$

$$R_{2} = \left[R^{2} + K_{12} - uK_{9} - vK_{10} - wK_{11}\right]^{1/2}$$

$$P = \frac{K_{16} + 2uK_{13} + 2vK_{14} + 2wK_{15}}{R_{0} + R_{3}}$$

Doppler Radar (Cont'd)

$$q = \frac{K_{20} + 2uK_{17} + 2vK_{18} + 2vK_{19}}{R_0 + R_2}$$

$$\dot{R}_{o} = \frac{\dot{u}(u - K_{24}) + \dot{v}(v - K_{25}) + \dot{w}(w - K_{26})}{R_{o}}$$

$$\dot{p} = \frac{\dot{u}K_{13} + \dot{v}K_{14} + \dot{w}K_{15} - p \dot{R}_{0}}{R_{1}}$$

$$\dot{q} = \frac{\dot{u}K_{17} + \dot{v}K_{18} + \dot{w}K_{19} - q\dot{R}_{0}}{R_{2}}$$

# Look Angles

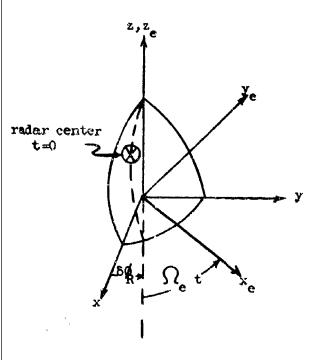
Missile look angles from radar site are

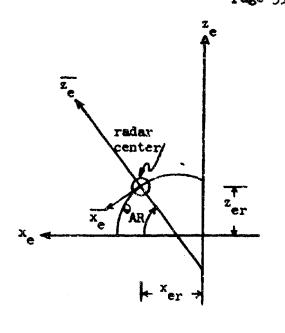
RIAL = 
$$\cos^{-1}\left[\frac{(x - x_r) \dot{\xi}_x + (y - y_r) \dot{\xi}_y + (z - z_r) \dot{\xi}_z}{\left[(x - x_r)^2 + (y - y_r)^2 + (z - z_r)^2\right]^{1/2}}\right]$$

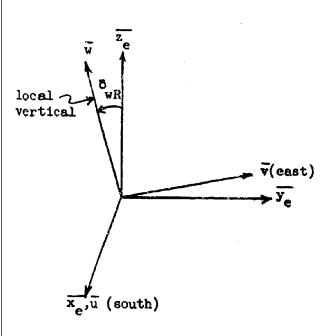
RIA2 = 
$$\tan^{-1} \left[ \frac{(x - x_r) \zeta_x + (y - y_r) \zeta_y + (z - z_r) \zeta_z}{(x - x_r) \eta_x + (y - y_r) \eta_y + (z - z_r) \eta_z} \right]$$

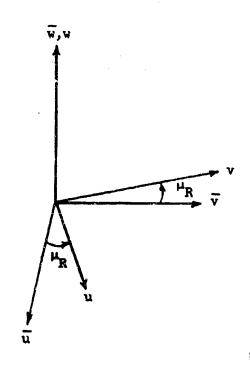
rad

lo: ve:



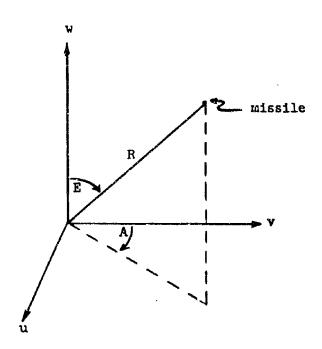






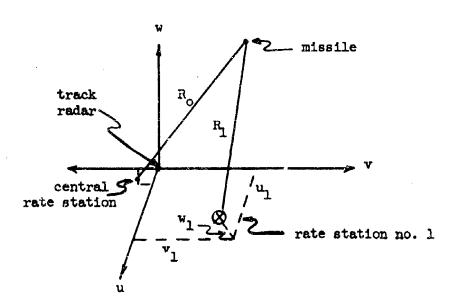
# COORDINATE TRANSFORMATIONS

FIGURE 11



SITE - FIXED RECTANGULAR TO SPHERICAL CONVERSION

FIGURE 12



DOPPLER RATE STATION SETUP

FIGURE 13

# Look Angles (Cont'd)

where

$$x_r = x_{er} \cos(\Omega_e t + \phi_r)$$

$$y_r = x_{er} \sin(\Omega_e t + \phi_r)$$

$$z_r = z_{er}$$

Pitch and yaw look angles

If RIA1 = 
$$90^\circ$$
 or  $270^\circ$ 

$$LAY = 90^{\circ}$$
 with opposite sign of RIAZ

If 
$$|RLA2| < 90^{\circ}$$
;  $LAP = -90^{\circ}$ 

If 
$$|RLA2| \ge 90^\circ$$
;  $LAP = 90^\circ$ 

If RIAL  $\neq 90^{\circ}$  or  $270^{\circ}$ 

$$tan RIAl = \frac{sinRIAl}{cosRIAl}$$

$$\overline{\text{LAP}} = \tan^{-1} \left[ -(\cos \text{RLA2})(\tan \text{RLA1}) \right]$$

$$IAY = \overline{IAY}$$

$$LAY = \overline{LAY} + 180^{\circ}$$

### Radar Conversions

# Look Angles (Cont'd)

If RIA2 > 
$$-90^{\circ}$$

If RIA2  $\leq 0^{\circ}$ 

IAP =  $-\left|\overline{\text{IAP}}\right|$ 

IAY =  $\overline{\text{IAY}}$  +  $180^{\circ}$ 

If RIA2 >  $0^{\circ}$ 

If RIA2  $\leq 90^{\circ}$ 

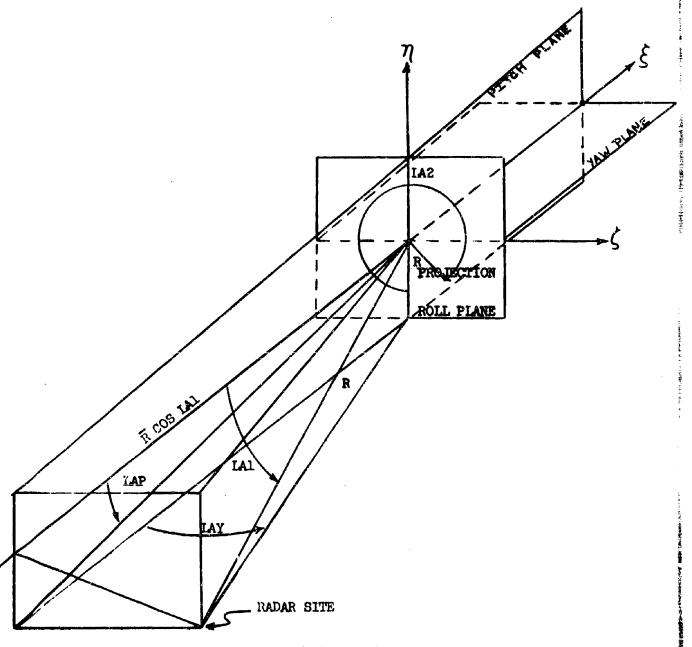
IAP =  $-\left|\overline{\text{IAP}}\right|$ 

IAY =  $\overline{\text{IAY}}$  -  $180^{\circ}$ 

If RIA2 >  $90^{\circ}$ 

IAP =  $\left|\overline{\text{IAP}}\right|$ 

IAY =  $\overline{\text{IAY}}$  -  $180^{\circ}$ 



and the second second

RADAR LOOK ANGLES

FIGURE 14

# Launch Azimuth Estimation

$$T_{\mathbf{F_i}} = \begin{cases} 1800, i = 0 \\ \kappa_0 + \kappa_1 e_{i-1} + \kappa_2 e_{i-1}^2, i > 0 \end{cases}$$

$$\theta_{i} = \cos^{-1} \left[ \sin \theta_{CL} \sin \theta_{CT} + \cos \theta_{CL} \cos \theta_{CT} \cos (\phi_{T} - \phi_{L} + \Omega_{e^{T}F_{i}}) \right]$$

Iterate until  $\left| e_{N} - e_{N-1} \right| < .0003$  radians

$$\bullet_{j} = \sin^{-1} \left[ \frac{\cos \theta_{N} \sin(\phi_{T} - \phi_{L} + \Omega_{e} T_{F_{N}})}{\sin \theta_{N}} \right]$$

$$v_{L} = -(x_{o}^{2} + y_{o}^{2})^{1/2} \Omega_{e} \cos a_{j}$$

$$v_d = (x_0^2 + y_0^2)^{1/2} \Omega_{e \text{ sina} j}$$

$$v_r = -v_d + KSTO + KSTI(\theta_N) + KST2(\theta_N)^2$$

$$T_p = K2STO + K2ST1(v_r) + K2ST2(v_r)^2$$

$$AZ_{L} = a_{j} + \frac{\frac{v_{L}}{v_{r}}}{1 - \frac{T_{p}}{3T_{F_{N}}}}$$

I.a

ξ

la

# Launch Azimuth Estimation (Cont'd)

$$K_0 = 220.37$$

$$K_1 = 20.621$$

$$K_2 = 0.$$

$$KSTO = 11137.6$$

$$KST2 = -.848166$$

$$K2STO = 141.675$$

$$K2ST1 = .0062625$$

$$K2ST2 = 0.$$

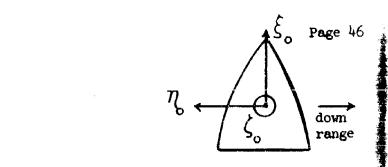
### Launch Attitude

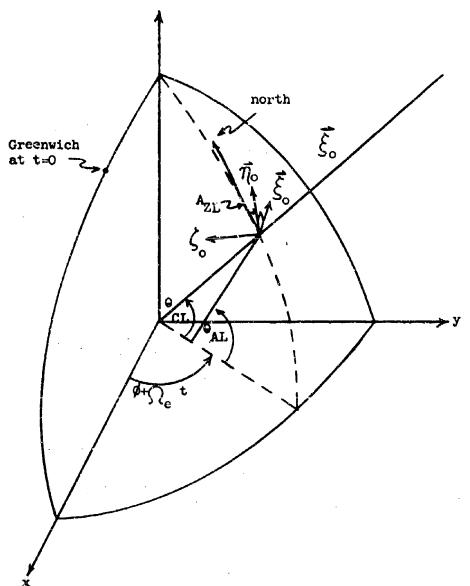
To achieve launch vertical the  $\xi, \bar{\eta}, \bar{\zeta}$  system is oriented such that  $\bar{\xi}$  points along the x direction,  $\bar{\eta}$  along the z direction, and  $\bar{\zeta}$  in the negative y direction. The  $\bar{\xi}, \bar{\eta}, \bar{\zeta}$  system is rotated through longitude of launch, astronomic latitude of launch, and the supplement of launch azimuth. (Fig. 15)

$$\vec{\xi} = \vec{\xi} \cos(-\phi_L - \Omega_e t) + \vec{\zeta} \sin(-\phi_L - \Omega_e t)$$

$$\vec{\eta} = \vec{\eta}$$

$$\vec{\zeta} = -\vec{\xi} \sin(-\phi_L - \Omega_e t) + \vec{\zeta} \cos(-\phi_L - \Omega_e t)$$





INITIAL ATTITUDE VECTORS

FIGURE 15

# Leunch Attitude (Cont'd)

then

$$\vec{\xi}' = \vec{\xi}''_{\cos(-\theta_{AL})} - \vec{\eta}''_{\sin(-\theta_{AL})}$$

$$\vec{\eta}' = \vec{\xi}''_{\sin(-\theta_{AL})} + \vec{\eta}''_{\cos(-\theta_{AL})}$$

$$\vec{\zeta}' = \vec{\zeta}''$$

finally

$$\overline{\eta}_{c} = \overline{\eta}^{\prime}_{\cos(AZ_{T} - 180)} - \overline{\zeta}^{i}_{\sin(AZ_{T} - 180)}$$

$$\vec{\zeta}_{0} = \vec{\eta}' \sin(AZ_{L} - 180) + \vec{\zeta}' \cos(AZ_{L} - 180)$$

Launch coordinate rotational quantity,  $\mu_{L}$  is estimated, if not specified, as

$$\mu_{L}$$
 = 180 - AZ<sub>L</sub>

<u>\$</u>\_\$

# Position and Velocity Components at Launch

Sea level radius of launch (Fig. 16) is

$$r_{SL_L} = \bar{A} \left[ \frac{1 + (1-e)^{\frac{1}{4}} \tan^2 \theta_{DL}}{1 + (1-e)^2 \tan^2 \theta_{DL}} \right]^{\frac{1}{2}}$$

# Position and Velocity Components at Launch (Cont'd)

Components of  $\overrightarrow{r}_{o}$  and  $\overrightarrow{v}_{o}$  are

$$x_o = \cos(\phi_L + \Omega_e t) \left[ r_{SL_L} \cos \theta_{CL} + (h_L + g_{SL}) \cos \theta_{DL} \right]$$

$$y_o = \sin(\phi_L + \Omega_e t) \left[ r_{SL_L} \cos \theta_{CL} + (h_L + g_{SL}) \cos \theta_{DL} \right]$$

$$z_o = r_{SL_L} sin\theta_{CL} + (h_L + g_{SL}) sin\theta_{DL}$$

$$\dot{x}_o = -y_o \Omega_e$$

$$\dot{y}_{o} = x_{o} \Omega_{e}$$

The simulation may be started at an arbitrary point in space with  $\vec{r}_o, \vec{v}_o, \vec{t}_o$ , and  $\vec{\eta}_o$  specified.

If the position and velocity components are given in a radar coordinate system, the following conversions are made:

18

LAUNCH COORDINATES ON REFERENCE ELLIPSOID

FIGURE 16

### Radar to Cartesian Coordinate Conversion

# Ellipsoid - Spherical Coordinates and Doppler Rates to Fixed Rectangular

u = R sinE sinA

v = R sinE cosA

 $W = R \cos E$ 

$$R_{0} = \left[R^{2} + K_{4} - uK_{1} - vK_{2} - wK_{3}\right]^{1/2}$$

$$R_{1} = \left[R^{2} + K_{8} - uK_{5} - vK_{6} - wK_{7}\right]^{1/2}$$

$$R_2 = \left[R^2 + K_{12} - uK_9 - vK_{10} - wK_{11}\right]^{1/2}$$

$$p = \frac{K_{16} + 2(uK_{13} + vK_{14} + wK_{15})}{R_0 + R_1}$$

$$q = \frac{K_{20} + 2(uK_{17} + vK_{18} + wK_{19})}{R_0 + R_2}$$

$$\dot{R}_{o} = \dot{R}$$

$$R_{o} \dot{p} = R_{p} \dot{p} + p(\dot{R}_{o} - \dot{p})$$

$$R_{0} \dot{q} = R_{0} \dot{q} + q(\dot{R}_{0} - \dot{q})$$

$$D = K_{23}(u-K_{24}) + K_{22}(v-K_{25}) + K_{21}(w-K_{26})$$

# Radar to Cartesian Coordinate Conversion (Cont'd)

$$\dot{\mathbf{u}} = \frac{R_{0}\dot{\mathbf{q}}}{D} \left[ K_{15}(\mathbf{v} - K_{25}) - K_{1\dot{\mathbf{q}}}(\mathbf{w} - K_{26}) \right] + \frac{R_{0}\dot{\mathbf{p}}}{D} \left[ K_{18}(\mathbf{w} - K_{26}) - K_{19}(\mathbf{v} - K_{25}) \right] + \frac{R_{0}\dot{\mathbf{R}}_{0}}{D} \cdot K_{23}$$

$$\dot{\mathbf{v}} = \frac{\frac{R_{o}\dot{\mathbf{q}}}{D} \left[ K_{13} (\mathbf{w} - K_{26}) - K_{15} (\mathbf{u} - K_{24}) \right] + \frac{\frac{R_{o}\dot{\mathbf{p}}}{D} \left[ K_{19} (\mathbf{u} - K_{24}) - K_{17} (\mathbf{w} - K_{26}) \right] + \frac{\frac{R_{o}\dot{\mathbf{R}}_{o}}{D}}{D} \cdot K_{22}$$

$$\dot{\mathbf{v}} = \frac{\frac{R_{o}\dot{\mathbf{q}}}{D} \left[ K_{1\dot{\mathbf{q}}} (\mathbf{u} - K_{2\dot{\mathbf{q}}}) - K_{1\dot{3}} (\mathbf{v} - K_{2\dot{5}}) \right] + \frac{\frac{R_{o}\dot{\mathbf{p}}}{D} \left[ K_{17} (\mathbf{v} - K_{2\dot{5}}) - K_{18} (\mathbf{u} - K_{2\dot{\mathbf{q}}}) \right] + \frac{\frac{R_{o}\dot{R}_{o}}{D}}{D} \cdot K_{2\dot{1}}$$

Doppler radar coefficients (see Fig. 13).

$$K_1 = 2(u_0)$$

$$K_2 = 2(v_0)$$

$$K_3 = 2(w_0)$$

$$K_{l_4} = (u_0)^2 + (v_0)^2 = d_0^2$$

$$K_5 = 2(u_1)$$

$$K_6 = 2(v_1)$$

$$\mathbb{K}_7 = 2(\mathbf{w}_1)$$

$$K_8 = (u_1)^2 + (v_1)^2 + (w_1)^2 = d_1^2$$

# Radar to Cartesian Coordinate Conversion (Cont'd)

$$K_9 = 2(u_2)$$
 $K_{10} = 2(v_2)$ 
 $K_{11} = 2(w_2)$ 
 $K_{12} = (u_2)^2 + (v_2)^2 + (w_2)^2 = d_2^2$ 

$$K_{13} = u_1 - u_0$$
 $K_{14} = v_1 - v_0$ 
 $K_{15} = w_1 - w_0$ 
 $K_{16} = d_0^2 - d_1^2$ 

$$K_{17} = w_2 - w_0$$
 $K_{18} = v_2 - v_0$ 
 $K_{19} = w_2 - w_0$ 
 $K_{20} = d_0^2 - d_2^2$ 

$$K_{21} = K_{13}K_{18} - K_{17}K_{14}$$
 $K_{22} = K_{15}K_{17} - K_{19}K_{13}$ 
 $K_{23} = K_{14}K_{19} - K_{18}K_{15}$ 

$$K_{24} = u_0$$

$$K_{25} = v_0$$

$$K_{26} = w_0$$

## Radar to Cartesian Coordinate Conversion (Cont'd)

Spherical coordinates and rates to ellipsoid-fixed rectangular coordinates

u = R sinE sinA

v = R sinE cosA

 $W = R \cos E$ 

 $\hat{\mathbf{u}} = \hat{\mathbf{R}} \sin \mathbf{E} \sin \mathbf{A} + \hat{\mathbf{R}} \hat{\mathbf{E}} \cos \mathbf{E} \sin \mathbf{A} + \hat{\mathbf{I}} \hat{\mathbf{A}} \sin \mathbf{E} \cos \mathbf{A}$ 

 $\dot{\mathbf{v}} = \dot{\mathbf{R}}$  sine  $\cos \mathbf{A} + \mathbf{R}\dot{\mathbf{E}}$   $\cos \mathbf{E}$   $\cos \mathbf{A} - \mathbf{R}\dot{\mathbf{A}}$  sine  $\sin \mathbf{A}$ 

 $\dot{\mathbf{w}} = \dot{\mathbf{R}} \cos \mathbf{E} - \mathbf{R} \dot{\mathbf{E}} \sin \mathbf{E}$ 

### Site Fixed Cartesian Coordinate and Rate Transformation to Inertial System

$$\begin{bmatrix} \dot{\xi} \\ \eta \\ \dot{\zeta} \end{bmatrix} = \begin{bmatrix} \cos \mu_{R} & -\sin \mu_{R} & 0 \\ \sin \mu_{R} & \cos \mu_{R} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} u \\ v \\ w \end{bmatrix}$$

$$\begin{bmatrix} \dot{\xi} \\ \dot{\eta} \\ \dot{\zeta} \end{bmatrix} = \begin{bmatrix} \cos \mu_{R} & -\sin \mu_{R} & 0 \\ \sin \mu_{R} & \cos \mu_{R} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \dot{u} \\ \dot{v} \\ \dot{w} \end{bmatrix}$$

$$\alpha = \xi$$

$$\beta = \eta - \delta_{W_{R}} \cdot \zeta \qquad \beta = \dot{\eta} - \delta_{W_{R}} \cdot \dot{\zeta}$$

$$\gamma = \delta_{W_{R}} \cdot \dot{\eta} + \dot{\zeta} \qquad \dot{\tau} = \delta_{W_{R}} \cdot \dot{\eta} + \dot{\zeta}$$

### Radar to Cartesian Coordinate Conversion (Cont'd)

$$\begin{bmatrix} \mathbf{x}_{e} \\ \mathbf{y}_{e} \\ \mathbf{z}_{e} \end{bmatrix} = \begin{bmatrix} \sin \theta_{\mathrm{DR}} & 0 & \cos \theta_{\mathrm{DR}} \\ 0 & 1 & 0 \\ -\cos \theta_{\mathrm{DR}} & 0 & \sin \theta_{\mathrm{DR}} \end{bmatrix} \begin{bmatrix} \alpha \\ \beta \\ \gamma \end{bmatrix} + \begin{bmatrix} \mathbf{x}_{er} \\ 0 \\ \mathbf{z}_{er} \end{bmatrix}$$

$$\begin{bmatrix} \dot{\mathbf{x}}_{e} \\ \dot{\mathbf{y}}_{e} \\ \dot{\mathbf{z}}_{e} \end{bmatrix} = \begin{bmatrix} \cos \theta_{\mathrm{DR}} & 0 & \sin \theta_{\mathrm{DR}} \\ 0 & 1 & 0 \\ -\sin \theta_{\mathrm{DR}} & 0 & \cos \theta_{\mathrm{DR}} \end{bmatrix} \begin{bmatrix} \dot{\alpha} \\ \dot{\beta} \\ \dot{\gamma} \end{bmatrix}$$

$$\vec{\mathbf{x}} = \begin{bmatrix} \cos(\Omega_{e} t + \delta \phi_{R} & -\sin(\Omega_{e} t + \delta \phi_{R}) & 0 \\ \sin(\Omega_{e} t + \delta \phi_{R} & \cos(\Omega_{e} t + \delta \phi_{R}) & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \mathbf{x}_{e} \\ \mathbf{y}_{e} \\ \mathbf{z}_{e} \end{bmatrix}$$

$$\vec{\mathbf{v}} = \begin{bmatrix} \cos(\Omega_{e} t + \delta \phi_{R}) & -\sin(\Omega_{e} t + \delta \phi_{R}) & 0 \\ \sin(\Omega_{e} t + \delta \phi_{R}) & \cos(\Omega_{e} t + \delta \phi_{R}) & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \dot{\mathbf{x}}_{e} \\ \dot{\mathbf{y}}_{e} \\ \dot{\mathbf{z}}_{e} \end{bmatrix} + \Omega_{e} \begin{bmatrix} -\mathbf{y} \\ \mathbf{x} \\ 0 \end{bmatrix}$$

# Attitude from Angles of Attack

The following conversions are made to initialize the attitude of the missile when  $\alpha_p$  and  $\alpha_v$  are known:

# Attitude from Angles of Attack (Cont'd)

$$\vec{v}_{a} = \vec{v} + \Omega_{e} \quad \begin{bmatrix} y \\ -x \\ 0 \end{bmatrix}$$

$$\vec{\xi}_{o}' = \frac{\vec{v}_{a}}{|\vec{v}_{a}|}$$

$$\cos \theta = \frac{1}{|\vec{r}|} \left[ x \cdot \xi_{x_0}' + y \cdot \xi_{y_0}' + z \cdot \xi_{z_0}' \right]$$

$$\sin\beta = \left[1 - \cos^2\beta\right]^{1/2}$$

$$\widehat{\widehat{\eta}}_{o}' = \left[ \frac{\widehat{r}}{|\widehat{r}|} - \cos\beta \cdot \widehat{\xi}_{o}' \right] \frac{1}{\sin\beta}$$

$$\vec{\zeta}' = \vec{\zeta}' \times \vec{\eta}'$$

$$\vec{\xi}_{o} = \cos \alpha_{y} \left[ \cos \alpha_{p} \cdot \vec{\xi} - \sin \alpha_{p} \cdot \vec{\eta} \right] + \sin \alpha_{y} \cdot \vec{\xi}$$

$$\vec{\eta}_{o} = \sin \alpha_{p} \cdot \vec{\xi}' + \cos \alpha_{p} \vec{\eta}'$$

$$\vec{\xi} = \vec{\xi} \times \hat{\eta}_{\circ}$$

### Tie Down

Derivatives of equations of motion to maintain vertical attitude during launch hold down are computed from

$$\omega_{\mathbf{r}} = -(\hat{\boldsymbol{\xi}}_{\mathbf{z}}) \cdot \hat{\boldsymbol{\Omega}}_{\mathbf{e}}$$

$$\omega_{\mathbf{p}} = -(\hat{\boldsymbol{\zeta}}_{\mathbf{z}}) \cdot \hat{\boldsymbol{\Omega}}_{\mathbf{e}}$$

$$\omega_{\mathbf{y}} = -(\hat{\boldsymbol{\eta}}_{\mathbf{z}}) \cdot \hat{\boldsymbol{\Omega}}_{\mathbf{e}}$$

$$\vec{\mathbf{v}} = -(\hat{\boldsymbol{\Omega}}_{\mathbf{e}})^{2} \cdot \begin{bmatrix} \mathbf{x} \\ \mathbf{y} \\ \mathbf{0} \end{bmatrix}$$

APPENDICES

# APPENDIX A

# Symbol Definitions

Ã	Semi-major axis of reference ellipsoid.
$\mathbf{A_{Z}}_{\mathrm{L}}$	Azimuth of launch, measured clockwise from north.
A <sub>Z</sub> L	Wind azimuth, measured clockwise from north.
B	Semi-minor axis of reference ellipsoid.
c	Speed of sound.
$\mathbf{c}_{_{\mathrm{D}}}$	Tabular drag coefficient.
e	Ellipticity of reference ellipsoid = $\frac{\overline{A}-\overline{B}}{\overline{A}}$ .
•	Eccentricity = $\frac{\overline{C}}{\overline{A}}$ , where $\overline{C}$ is the distance from the center
	of the ellipsoid to a focus along the major axis.
G	Gravitational Acceleration.
ø	Longitude.
θ <sub>A</sub>	Astronomic latitude, angle between the equitorial plane and the local direction of gravity.
<sup>e</sup> c	Geocentric latitude, angle between the equatorial plane and T.
$\Theta_{\mathrm{D}}$	Geodetic latitude, angle between the equatorial plane and the normal to the reference ellipsoid.
K <sub>F</sub>	Multiplier for tabular thrust $F_{o}$ .
K <sub>p</sub>	Wind perturbation multiplier.
К <sub>М</sub>	Multiplier for tabular mass flow rate $\dot{M}_{\rm O}$ .
K <sub>w</sub> pt K <sub>wrt</sub>	Multiplier for tabular pitch rate.
K <sub>wrt</sub>	Multiplier for tabular roll rate.

Multiplier for tabular yaw rate.

# APPENDIX A (Cont'd)

м	Mass of missile.
M <sub>I</sub>	Reference initial mass for ablation computation.
$N_{\mathbf{m}}$	Mach number.
P	Atmospheric pressure.
Q	Dynamic pressure.
ρ	Atmospheric density.
r	Missile position vector whose components are x,y,z.
$\mathbf{r}_{o}$	Average radius of geoid (spherical approximation).
S	Missile cross-sectional area.
T	Atmospheric temperature.
ø	Longitude.
V <sub>w</sub>	Tabular wind velocity.
<b>x</b> , <b>y</b> ,ż	Components of missile's velocity vector $\overrightarrow{V}$ .
$\Omega_{ m e}$	Reference ellipsoid's rotation rate.
wp,wy,wr	Angular rotational rate about respectively pitch, yaw and roll axes.
p <sub>c</sub> , y <sub>c</sub> , w <sub>r</sub>	Respectively pitch, yaw and roll rates from the guidance program.
wpd,wyd,wrd	Respectively pitch, yaw and roll drifts.

# APPENDIX B Iteration

Two general schemes of iteration are used. The first, called (for lack of a better name) "paired function", requires that each parameter to be determined be paired with a function to be constrained. The second form of iteration, termed "matrix", does not require pairing of the functions and parameters but does not allow minimizing or maximizing of functions.

In either type of iteration, control of the portions of a trajectory over which iteration is to take place is provided. The control system allows "nesting" of "paired function" sets within other "paired function" or "matrix" sets. This capability allows combination of the strong points of the two general schemes according to the needs of the individual problem.

### Paired Functions

After an initial trajectory is integrated, a perturbation is added to the parameter in question and trather trajectory is integrated. Inverse La Grangian interpolation or extrapolation is then used to estimate a new value for the parameter. (The desired value of the function is used as argument) The trajectory is again integrated and the results enable the use of a higher order polynomial in the interpolation or extrapolation procedure. The process is used repeatedly until convergence. The polynomial order may be limited by input, or in any case at order 5.

If minimization or maximization is called for, a second perturbation is added to the parameter and a third trajectory is produced. La Grangian interpolation is then used in a "step-and-bisect" iterative procedure which determines minimum or maximum for the polynomial fitting the available points. If no minimum or maximum is available, another perturbation is added to the parameter. In either case another trajectory is produced and the new point is used to enable a higher order polynomial to be used in the "step-and-bisect" solution. Again the La Grangian polynomial order may be limited by input, or at order 5.

# Paired Functions (cont'd)

The "step-and-bisect" process is said to converge when the improvement in the function is less than one half the (input) accuracy requirement. The function maximization (minimization) is complete when the difference between output of the "step-and-bisect" procedure and the previous value of the function is less than the (input) accuracy requirement.

# MATRIX Iteration

The Newton-Raphson method for iterative linear differential correction may be stated in its matrix form:

$$\begin{array}{c|c} \vdots & \ddots & \frac{\mathcal{Y}_{x^{U}}}{\mathcal{Y}_{x^{T}}} \\ \hline & \frac{\mathcal{Y}_{x^{T}}}{\mathcal{Y}_{x^{T}}} & \frac{\mathcal{Y}_{x^{S}}}{\mathcal{Y}_{x^{T}}} \end{array}$$

Where  $\Delta Y$  consists of the differences from desired conditions, and  $\Delta X$  is the suggested changes to the parameters in question. The rettial derivative elements  $\Delta Y_1$  may be replaced by their finite difference equivalents,  $\Delta Y_1$ . In the N-STAGE program, in order to enable the use of the  $\Delta X_1$ 

Generalized Secont Method, the resultant equation is re-written as two equations in the equivalent form:

$$\begin{bmatrix} \Delta Y \\ -\Delta Y \end{bmatrix} = \begin{bmatrix} \Delta^{Y}_{1,1} & \Delta^{Y}_{1,2} \\ \Delta Y_{2,1} & \vdots \\ \vdots & \cdots & \Delta^{Y}_{m,n} \end{bmatrix}$$

(3) 
$$\begin{bmatrix} \Delta x \\ 0 \end{bmatrix} = \begin{bmatrix} \Delta x_1 & 0 \\ 0 & \Delta x_2 \\ \vdots & \Delta x_n \end{bmatrix} \begin{bmatrix} p \\ -1 & -1 \\ 0 & -1 \end{bmatrix}$$

Where  $\triangle Y_{i,j}$  indicates the difference in  $Y_i$  produced by a change  $\triangle X_j$  in  $X_j$ . The elements  $\triangle Y_{i,j}$  are obtained by repeatedly integrating the trajectory with ceach of the parameters in question perturbed in its turn. The elements  $\triangle X_j$  of the diagonal matrix (3), are the perturbation values. The elements of  $\triangle Y_j$  and  $\triangle Y_i$ , are divided by normalizing elements  $N_i$  (which are input quantities) in order to scale, or establish dimensional homogeneity.

The first equation is solved for [P] in a "least squares"\* sense, which allows the general case, m/n. The multiplication implied in the second equation is then performed, yielding the set  $[\Delta x]$ , the elements of which are added to the appropriate parameters.

After a solution of the above sort is accomplished, and the trajectory representing the suggested parameters integrated and found to be still unconverged, the Generalized Secant Method may be utilized to improve the convergence without integrating the number of trajectories (n), necessary to recompute the matrices for equations (2) and (3). In this method a new column, consisting of the negative of the normalized (scaled) deviations from conditions at the previous nominal trajectory integration, is appended to each matrix as the first column, and is added to each of the remaining columns, producing matrices with one more column than the previous matrices.

$$\begin{bmatrix} -\Delta Y \\ -\Delta Y \end{bmatrix} = \begin{bmatrix} -\Delta Y_{1,0} & \Delta Y_{1,1} - \Delta Y_{1,0} & \cdots \\ -\Delta Y_{2,0} & \cdots & \Delta Y_{m,n} - \Delta Y_{m,o} \end{bmatrix}$$

<sup>\*</sup>See RWLS4F, General Linear Equation Solver by D. D. Morrison, Sept. 24, 1959 (CDRC minister subroutine).

(5) 
$$\begin{bmatrix} \triangle x \end{bmatrix} = \begin{bmatrix} -\triangle x_{1,0} & \triangle x_{1,1} - \triangle x_{1,0} & \ddots & \\ -\triangle x_{2,0} & -\triangle x_{2,0} & \ddots & \\ \vdots & \ddots & \triangle x_{n,0} \end{bmatrix}$$

The process may be used repeatedly, each time defining the previous trajectory to be the previous nominal trajectory.

The last column may be deleted after any number of additional columns have been accumulated.

A satisfactory rationale for deletion or non-deletion of the last column is not presently available, but experience indicates that the last column should be kept if the number of functions is approximately equal to the number of parameters (m=n), and deleted when the number of functions is markedly greater than the number of parameters (m=n). Tests with m and n initially equal to 3 have shown a 20% increase in efficiency when the matrices are allowed to expand indefinitely.

In particular, when inaccuracies due to system non-linearities (round-off error accumulations, integration truncation error, staging condition annomalies, etc.) are on the order of the iteration convergence accuracies required, it has been found necessary to allow the matrices to expand in order to gain convergence.

An additional advantage of the Generalized Secant Method is that perturbation delta magnitudes essume a less dominant role in convergence efficiency. In many cases convergence is worsened by a Newton-Raphson solution because of erroneous partial derivative information due to improper perturbation sizes. In such cases it is not uncommon to find that the Generalized Secant Method can proceed to convergence.

The Generalized Secant Method is used repeatedly, until the gain in one step is less than the input quantity SCGN. (The gain is measured as the ratio of the summations of the squares of the elements of the normalized error vector  $\Delta Y$ , for two successive solution passes.) Thus, if any solution pass (whether the solution comes

from Newton-Raphson or Generalized Secant Method) is not materially better than the previous solution or nominal pass, the program recomputes the matrices by perturbation methods. Note that if SCGN is a very large number, the Generalized Secant Nethod is not attempted.

Eigenvalues associated with the solution of equation (2) or (4) are printed following each evaluation. These Eigenvalues are of use only if it is understood that they are for the matrix transpose multiplied by the matrix, and it should be remembered that the elements are scaled. The ratio of the largest Eigenvalue to the smallest (known as the condition number) is sometimes a clue to the degree of linear dependence in the matrix. A possibly confusing issue is that the overall perturbation in the column associated with a particular Eigenvalue also affects the condition number. Zero Eigenvalues should not be found for a Newton-Raphson solution, but are often found for a Generalized Secant Method solution. In fact, usually only (n) Eigenvalues are large enough to have any effect, even though the matrix has been allowed to expand considerably larger than the original (n) columns.

In any case, the Eigenvalues are associated with input parameters in reverse order to their printing. That is, the last Eigenvalue is associated with the first iteration variable.

- A matrix iteration is determined to have converged on either of two criteria.
  - 1. The error in each of the functions is less than the corresponding accuracy requirement.
  - 2. The change in the error (for each of the functions which fails the first criterion) is less than one half the corresponding accuracy requirement.

    The change is measured from the previous solution or the nominal trajectory.

#### APPENDIX C

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